CSE 252B: Computer Vision II

Lecturer: Serge Belongie Scribe: Smriti Yamini

LECTURE 7 Uncalibrated Epipolar Geometry

7.1. Uncalibrated camera or distorted space?

So far we have been assuming K = I, which corresponds to the calibrated case. What happens when $K \neq I$? In 1992, Faugeras asked the question "What can be seen in 3-D with an uncalibrated stereo rig?" That is, what can we determine about the 3-D structure of the scene and the pose of the camera in the uncalibrated case? Hartley also posed the same question. The answer to the question is this: you can recover the structure of the scene in 3D (and the camera pose) up to a projective transformation.

There are in fact two equivalent ways to look at the problem setup:

- an uncalibrated camera moving in rectified space, or
- a calibrated camera moving in distorted space.

An uncalibrated camera with calibration matrix K, viewing points in a calibrated (Euclidean) world and moving with parameters (R, \mathbf{T}) is equivalent to a calibrated camera viewing points in distorted space moving with parameters $(KRK^{-1}, K\mathbf{T})$. This is illustrated in Figure 1. This distorted

¹Department of Computer Science and Engineering, University of California, San Diego.

April 19, 2004

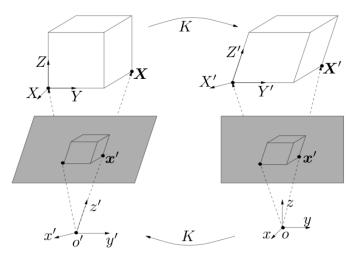


Figure 1. MaSKS Figure 6.4.

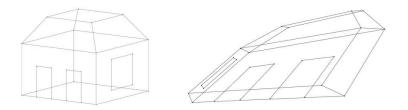


Figure 2. The Euclidean (left) and projective (right) structure of a house.

space is governed by the inner product

$$\langle \boldsymbol{u}, \boldsymbol{v} \rangle_S = \boldsymbol{u}^\top S \boldsymbol{v}, \text{ with } S = K^{-T} K^{-1} = (K K^\top)^{-1}$$

We call S the *metric* of the space. In the Euclidean case, S = I and $\langle \boldsymbol{u}, \boldsymbol{v} \rangle_S = \boldsymbol{u}^\top \boldsymbol{v}$. Figure 2 shows the difference between the Euclidean structure and the projective structure of a 3D object.

Recall the structure of the matrix K:

$$K = \begin{bmatrix} fs_x & s_\theta & o_x \\ 0 & fs_y & o_y \\ 0 & 0 & 1 \end{bmatrix}$$

Its purpose is to map metric coordinates (unit of metres) into image coordinates (unit of pixels). We use a prime to denote pixel coordinates:

$$\boldsymbol{x} = K^{-1} \boldsymbol{x}'$$

Applying the rotation matrix and the translation vector to some point X_0 in Euclidean space, we get

$$\boldsymbol{X} = R\boldsymbol{X}_0 + \boldsymbol{T}$$

In the uncalibrated camera frame we have

$$K \boldsymbol{X} = K R \boldsymbol{X}_0 + K \boldsymbol{T}$$
 or $\boldsymbol{X}' = K R K^{-1} \boldsymbol{X}'_0 + \boldsymbol{T}'$

where $\mathbf{X}' = K\mathbf{X}, \, \mathbf{T}' = K\mathbf{T}$ and $\mathbf{X}'_0 = K\mathbf{X}_0$.

Applying the image formation model using homogeneous coordinates, we get

$$\lambda \boldsymbol{x} = K \Pi_0 g \boldsymbol{X}_0$$

= $K \begin{bmatrix} I & \mathbf{0} \end{bmatrix} \begin{bmatrix} R & T \\ \mathbf{0}^\top & 1 \end{bmatrix} \boldsymbol{X}_0$
= $K R \boldsymbol{X}_0 + K \boldsymbol{T}$
= $K R K^{-1} \boldsymbol{X}'_0 + K \boldsymbol{T}$
= $\Pi_0 g' \boldsymbol{X}'_0$

where $g' = \begin{bmatrix} KRK^{-1} & \mathbf{T}' \\ \mathbf{0}^{\top} & 1 \end{bmatrix} \in \mathbb{R}^{4 \times 4}$ is the distorted rigid transformation.

Summarizing, an uncalibrated camera moving in the calibrated space $(\lambda \boldsymbol{x'} = K \Pi_0 g \boldsymbol{X}_0)$ is equivalent to a calibrated camera moving in a distorted space $(\lambda \boldsymbol{x'} = \Pi_0 g' \boldsymbol{X'}_0)$.

7.2. Epipolar Constraint

Recall the form of the epipolar constraint in the calibrated case:

$$\boldsymbol{x}_2^{\top} E \boldsymbol{x}_1 = 0$$

By direct substitution of the relation between metric and pixel coordinates $\boldsymbol{x} = K^{-1} \boldsymbol{x}'$, we get

$$\boldsymbol{x}_{2}^{\prime \top} \boldsymbol{K}^{-\top} \widehat{\boldsymbol{T}} \boldsymbol{R} \boldsymbol{K}^{-1} \boldsymbol{x}_{1}^{\prime} = \boldsymbol{0}$$

The matrix in the middle is known as the Fundamental matrix,

$$F = K^{-\top} \widehat{T} R K^{-1} = K^{-\top} E K^{-1}$$

Note that F reduces to the essential matrix when K = I.

7.2.1. Coplanarity constraint

We can also examine uncalibrated epipolar geometry in terms of the coplanarity constraint. The three vectors $\mathbf{x}'_2, \mathbf{T}' = K\mathbf{T}$ and $KR\mathbf{x}_1 = KRK^{-1}\mathbf{x}'_1$ in Figure 3 are coplanar. Hence their scalar triple product is 0:

$$\boldsymbol{x}_{2}^{\prime \top} \widehat{T}^{\prime} K R K^{-1} \boldsymbol{x}_{1}^{\prime} = 0$$

The matrix in the middle looks slightly different than the previous expression for F, but we will see shortly that they are equivalent.

7.2.2. Algebraic derivation

Start with

$$\lambda_2 \boldsymbol{x}_2 = R \lambda_1 \boldsymbol{x}_1 + \boldsymbol{T}$$

with $\lambda \boldsymbol{x} = \boldsymbol{X}$. Multiply both sides by K,

$$\lambda_2 K \boldsymbol{x}_2 = K R \lambda_1 \boldsymbol{x}_1 + K \boldsymbol{T}$$

or

$$\lambda_2 \boldsymbol{x}_2' = KRK^{-1}\lambda_1 \boldsymbol{x}_1' + \boldsymbol{T}'$$

Taking the dot product with $T' \times x'_2 = \widehat{T}' x'_2$ and dropping scalar factors, we get

$$\boldsymbol{x}_{2}^{\prime \top} \widehat{T}^{\prime} K R K^{-1} \boldsymbol{x}_{1}^{\prime} = 0$$

since the vector $\hat{T}' \boldsymbol{x}_2'$ is orthogonal to both \boldsymbol{T}' and \boldsymbol{x}_2' .

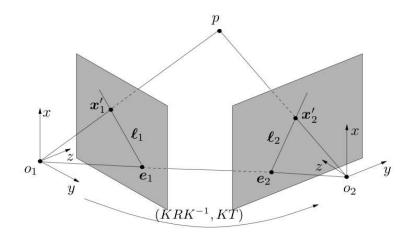


Figure 3. Uncalibrated epipolar geometry.

From MaSKS Lemma 5.4, we have the identity $K^{-T}\widehat{T}K^{-1} = \widehat{KT}$ when $\det(K) = +1$. We can therefore write

$$F = K^{-T} T R K^{-1} \quad \text{(uncalibrated camera in calibrated space)}$$
$$= K^{-T} T K^{-1} K R K^{-1}$$
$$= T' K R K^{-1} \quad \text{(calibrated camera in uncalibrated space)}$$

7.2.3. Epipolar Lines and Epipoles

The bilinear relationship

$$\boldsymbol{x}_2^{\prime \top} F \boldsymbol{x}_1^{\prime} = 0$$

transfers a point in view 1 to a line in view 2. Equivalently, we may write

$$\boldsymbol{x}_{2}^{\prime op} F \boldsymbol{x}_{1}^{\prime} = \boldsymbol{x}_{2}^{\prime op} \boldsymbol{l}_{2} = 0$$

where l_2 is the epipolar line in view 2. Similarly for l_1 , we have

$$\boldsymbol{l}_1^{ op} \boldsymbol{x}_1' = 0$$

Thus $\boldsymbol{l}_2 = F \boldsymbol{x}_1'$ and $\boldsymbol{l}_1 = F^\top \boldsymbol{x}_2'$.

As with E, we can get the epipoles from the left and right null space of F:

$$\boldsymbol{e}_2^{\mathsf{T}}F = \boldsymbol{0}, \quad F\boldsymbol{e}_1 = \boldsymbol{0}$$

from which it follows

$$\boldsymbol{e}_2 = K\boldsymbol{T} = \boldsymbol{T}', \quad \boldsymbol{e}_1 = KR^\top \boldsymbol{T}$$

7.3. Properties of F

Like the essential matrix, F has rank 2 since $\widehat{T'}$ is rank 2. The SVD of $F = U\Sigma V^{\top}$ has $\Sigma = \text{diag}\{\sigma_1, \sigma_2, 0\}$ with $\sigma_1 \geq \sigma_2$. (In the case of E, we had $\sigma_1 = \sigma_2$). This implies that any rank 2 matrix can be a fundamental matrix for some stereo rig.

We can apply the 8-point algorithm to estimate F using a design matrix $\chi \in \mathbb{R}^{n \times 9}$ with rows (carrier vectors) of the form

$$oldsymbol{a} = oldsymbol{x}_1' \otimes oldsymbol{x}_2'$$

(7.1)
$$\boldsymbol{a} = [x_1' x_2', x_1' y_2', x_1', y_1' x_2', y_1' y_2', y_1', x_2', y_2', 1]^\top \in \mathbb{R}^9$$

All variables are primed, as they are in pixel coordinates. If we assume the pixel coordinates are on the order of 10^2 , then we encounter a practical problem: the entries in \boldsymbol{a} range from 10^0 to 10^4 , which makes χ ill-conditioned. Hartley proposed a simple means of overcoming this problem.

7.4. Hartley Normalization

In Hartley normalization we rescale the data using two matrices H_i , i = 1, 2, so as to produce coordinates that make the design matrix well-conditioned. We choose H_i such that the normalized coordinates $\tilde{\boldsymbol{x}}_i = H_i \boldsymbol{x}'_i$ have zero mean and unit variance:

$$H_{i} = \begin{bmatrix} 1/\sigma_{x_{i}} & 0 & -\mu_{x_{i}}/\sigma_{x_{i}} \\ 0 & 1/\sigma_{y_{i}} & -\mu_{y_{i}}/\sigma_{y_{i}} \\ 0 & 0 & 1 \end{bmatrix}$$

In this expression, means and variances are given by:

$$\mu_{x_i} = \frac{1}{n} \sum_{j=1}^n x_i^j \qquad \sigma_{x_i}^2 = \frac{1}{n} \sum_{j=1}^n (x_i^j)^2 - \mu_{x_i}^2$$
$$\mu_{y_i} = \frac{1}{n} \sum_{j=1}^n y_i^j \qquad \sigma_{y_i}^2 = \frac{1}{n} \sum_{j=1}^n (y_i^j)^2 - \mu_{y_i}^2$$

Intuitively, H_i can be thought of as a guess at the calibration matrix, placing the centroid of the coordinates at the image center, assuming zero skew, and using the x and y variance to set the pixel aspect ratio.

After this transformation, we run the 8-point algorithm on \tilde{x}_i , i = 1, 2, to obtain the fundamental matrix \tilde{F} for the normalized data. Finally we obtain F by observing

$$\boldsymbol{x}_{2}^{\prime \top}F\boldsymbol{x}_{1}^{\prime} = \widetilde{\boldsymbol{x}}_{2}^{\top}\underbrace{\boldsymbol{H}_{2}^{-\top}F\boldsymbol{H}_{1}^{-1}}_{\widetilde{F}}\widetilde{\boldsymbol{x}}_{1} = 0$$

so $F = H_2^\top \widetilde{F} H_1$